# **Why do the second row transition metal atoms prefer 5s**<sup>1</sup>**4d**<sup>m + 1</sup> to 5s<sup>2</sup>**4d**<sup>m</sup>?

**Numerical Hartree-Fock studies** 

# Toshikatsu Koga<sup>1</sup>, Hidenori Aoki<sup>1</sup>, Hiroshi Tatewaki<sup>2</sup>

<sup>1</sup> Department of Applied Chemistry, Muroran Institute of Technology, Muroran, Hokkaido 050, Japan

2 Computation Center, Nagoya City University, Nagoya, Aichi 467, Japan

Received February 15, 1995/Final revision received March 30, 1995/Accepted March 31, 1995

**Abstract.** Numerical Hartree-Fock (NHF) calculations have been performed for 332 ground and low-lying excited states of the fifth period atoms  $R\bar{b}$  through Xe, with our special interest in the states arising from the  $5s^24d^m$ ,  $5s^14d^{m+1}$ , and  $5s^6$  $4d^{m+2}$  configurations of the second row transition metal atoms. Among various properties, orbital energies and mean values of r of the outermost orbitals of each symmetry are presented as well as total energies. It is discussed in some detail why the second row transition metal atoms have a tendency to prefer  $s^1 d^{m+1}$  as the ground configuration in contrast to the preferred  $s^2d^m$  configuration in the first row transition metal atoms. Our systematic NHF computations reported in this and the previous papers conclude that the Hartree-Fock method correctly predicts the experimental ground state of the atoms He through Xe with the sole exception for Zr.

**Key words:** Transition metal atoms **-** Ground and low-lying excited states **-**  Numerical Hartree-Fock approximation

# **1. Introduction**

Numerical Hartree-Fock (NHF) calculations have been performed for the ground states of many neutral atoms, cations [1-9], and anions [9]. In contrast to the ground state, few NHF results have been published [4, 5, 10-12] for excited states of atoms and atomic ions. Recently, we reported systematic NHF results for low-lying excited states of the atoms He to  $\bar{A}r$  [13], K to Kr [14], and cationic species from the lanthanide atoms [15]. In the present study, we consider the fifth period atoms Rb through Xe with the following configurations:



All the present NHF calculations were performed with an enhanced version of MCHF72 [16], and care was taken to ensure that the total energies are accurate to nine or more significant figures. Although we have calculated total 332 states arising from the above configurations, we give the information on the ground state, and the lowest and highest excited states of the respective configurations.

## **2. Results and discussion**

Table 1 summarizes the total energies, excitation energies, orbital energies  $\varepsilon_i$ , and mean electron-nucleus distances,  $\langle r \rangle_i$ , for the outermost orbitals of each symmetry  $i (= s, p, d)$ . All values are given in Hartree atomic units.

## *2.1. Atoms Rb and Sr*

In the Rb atom, the ground electron configuration is  $\lceil Kr \rceil 5s^1$ , where the 5s orbital is diffuse reflecting a large screening effect of the Kr core. As an angular momentum of the orbital increases, in which the excited electron moves, the diffuseness largely increases. In Sr, the 5p orbital of  $5s<sup>15</sup>p<sup>11</sup>P$  is fairly diffuse in comparison with that of  $5s<sup>15</sup>p<sup>13</sup>P$ . The broadness of the 5p atomic orbital (AO) in Sr<sup>1</sup>P arises from a repulsive exchange potential between the 5s and 5p electrons as in cases of Be  $2s<sup>1</sup>$  $2p^{1 \bar{1}}P$  [5, 10, 11, 17, 18], Mg  $3s^{1}3p^{1 \bar{1}}P$  [13], and Ca  $4s^{1}4p^{1 \bar{1}}P$  [14].

#### *2.2. Atoms Y through Cd*

In Fig. 1, we plot the total energy difference  $(\Delta E)$  between the lowest state arising from  $ns^1(n-1)d^{m+1}$  and that from  $ns^2(n-1)d^m$ :

$$
\Delta E = \text{TE}[ns^{1}(n-1)d^{m+1}] - \text{TE}[ns^{2}(n-1)d^{m}], \tag{1}
$$

where TE stands for the total energy and  $n$  is a principal quantum number of AOs;  $n = 4$  for the first row transition atoms and  $n = 5$  for the second row transition atoms. It is seen that for the first row transition metal atoms the lowest state from  $4s^23d^m$  is lower than that from  $4s^13d^{m+1}$  except for Cr and Cu. The ground state therefore arises from  $4s^2 3d^m$  except these two atoms. For the second row transition metal atoms, however, the lowest state from  $5s<sup>1</sup>3d<sup>m+1</sup>$  is lower than that from  $5s<sup>2</sup>4d<sup>m</sup>$  except for Y and Mn and the ground state has the configuration of  $5s<sup>1</sup>4d<sup>m+1</sup>$  except for Y, Tc, and Pd: the ground configuration for Pd is  $5s<sup>0</sup>4d<sup>10</sup>$ . The reason is discussed below for the different preference of the ground configurations between the first and second row transition metal atoms.

In Fig. 2, we give the  $d$  orbital energies of the lowest states arising from the  $4s^23d^m$ ,  $4s^13d^{m+1}$ ,  $5s^24d^m$ , and  $5s^14d^{m+1}$  configurations. It is interesting that the 4d orbital energies of  $5s<sup>1</sup>4d<sup>m+1</sup>$  are nearly equal to the 3d orbital energies of  $4s<sup>1</sup>3d<sup>m+1</sup>$ . Then the Koopmans' theorem suggests the similarity of the  $d$  ionization energies in the  $5s<sup>1</sup>4d<sup>m+1</sup>$  and  $4s<sup>1</sup>3d<sup>m+1</sup>$  configurations, but this is not true. We have calculated ionization energies (IEs) using *ASCF* method:

$$
IE = TE[ns1(n-1)dm] - TE[ns1(n-1)dm+1].
$$
 (2)

Results are shown in Fig. 3 where IEs for K, Ca, Rb, and Sr are defined as

$$
IE = TE[ns^{m-1}] - TE[ns^{m-1}(n-1)d^{1}] \quad (m = 1, 2). \tag{3}
$$







284





					s, p, and d orbitals	Orbital energies of outermost		s, p, and d orbitals	Mean value of r for outermost	
	Atom Configuration	State <sup>a</sup>	$-TE$	ДE	$-\varepsilon_s$	$\frac{d}{3}$	$ \varepsilon_d$	$\hat{\mathcal{L}}$	$\hat{\mathcal{L}}_p$	$\hat{\mathcal{L}}$
	$[Kr]$ 4d(10)5s(1)5p(4)		313.23804	1.24728	1,72771	132194	.73825	.36043	.88659	.01451
		$\mathbf{r}_z$	5313.01768	146764	1,53600	0.28104	.77696	.30138	.01120	.01353
	52 Te [Kr] $4d(10)5s(2)5p(4)$	Р,	5611.78406	00000	1.70056	135983	103829	1.22797	2.69137	0.96118
			5611.74140	04266	1.70488	0.34146	0.04529	1.22607	2.71099	0.96106
			611.67829	1.10577	1.71161	1.31471	1.05606	.22309	11211	1.96088
	[ $Kr$ ] $4d(10)5s(1)5p(5)$		6611.45555	1.32851	181616	1.35478	2.09107	2.20164	2.68072	195959
		ېم	5611.29499	1,48907	0.67144	0.33211	2.11393	2.17732	2.72833	0.95911
$\overline{53}$		$\mathbf{r}_2$	6917.98090	0.00000	1.82112	0.40318	2.40121	1.09443	2.50175	
	$[{\rm Kr}] \; 4d(10)5s(2)5p(5) \nonumber \\ [{\rm Kr}] \; 4d(10)5s(1)5p(6)$	S <sub>2</sub>	6917.56518	141572	0.90035	1.39993	2.45638	2.07051	2.49706	0.91303 0.91168
	54 Xe [Kr] 4d(10)5s(2)5p(6) [Kr] 4d(10)5s(2)5p(5)6s(1)		232.13836	000000	1,94441	1.45729	2.77788	96086		0.87045
			1231.82958	1,30878	1,12108 1,11531	1.61595	2.97178	,44840	2.33798	1.86956
			731.82424	31412		162092		15497	2.22915	1,86957
	$[Kr]$ 4d(10)5s(2)5p(5)6p(1)		231.80035	1.33801	.11896	1.09167	1,9997	93230	'.81434	1.86953
			231.78756	135080	.13432	1,07861	1.01497	93206	86498	1.86954
	Sd(1) $[Kr]$ 4d(10)5s(2)5p(5):	۹F	231.79117	134719	05329	1.58644	0.08438	94118	1.25416	5.47615
			231.76552	37284	.15107	1.67499	1.05658	.93178	1.22106	0.41388

<sup>&</sup>quot;The symbol in the brackets denotes the coupling in the 4d shell

286

Table 1. (Continued)



Fig. 1. Total energy difference between the lowest states from  $ns^2(n-1)d^m$  and  $ns^{1}(n-1)d^{m+1}$  configurations

Fig. 2. The d orbital enegies for  $ns^2(n-1)d^m$ and  $ns^1(n-1)d^{m+1}$  configurations

In Fig. 3, we find that the IEs for the second row transition metal atoms are always larger than those of the corresponding atoms in the first row transition series. The IE difference between the two series of the transition metal atoms increases as we move along the period. Thus, the electron in  $4d$  AO is stabler than that in the  $3d$ AO in the  $s<sup>1</sup>d<sup>m+1</sup>$  configuration. The sudden decrease of IE at Mn and Tc is due to the fact that all the spins of the outermost  $s$  and  $d$  electrons are parallel in the ionic states and the precorrelation effect is larger for the ions than for the corresponding neutral atoms.

We define a screening constant  $S$  by

$$
IE = (Z - S)^2 / 2n^2,
$$
 (4)

where Z stands for the nuclear charge. We consider S for 5s and 4d in the field generated by  $5s<sup>1</sup>4d<sup>m</sup>$  and S for 4s and 3d in the field of  $4s<sup>1</sup>3d<sup>m</sup>$ . The IE and S obtained from NHF and experiment  $[19]$  are given in Fig. 4 and Table 2, where the S values are subtracted by 18 and 36 for the first and second row transition metal atoms, respectively. We express the screening constant of  $4s$  in  $4s^23d^m$  as  $S_{4s}(4s^23d^m)$  or  $S_{4s}$  and that of 5s in  $5s^24d^m$  as  $S_{5s}(5s^24d^m)$  or  $S_{5s}$ . In Fig. 4, the screening constants  $S_{4s}$  and  $S_{5s}$  are almost parallel and increase linearly as the number of the *d* electrons increases. The smaller values of  $S_{ss}$ -36 than  $S_{4s}$ -18



Fig. 3. The d ionization energies for  $ns^{1}(n-1)d^{m+1}$  configurations calculated with ASCF

Fig. 4. Screening constants for d orbitals of  $ns^{1}(n-1)d^{m+1}$  and for s orbitals of  $ns^2(n-1)d^m$  calculated from the  $\triangle$ SCF ionization energies. The values 18 and 36 are subtracted, respectively, for the fourth- and fifth-period atoms

indicate the incomplete screening of the Kr core compared with that of the Ar core. The NHF screening constants for 4s and 5s are described by

$$
S_{4s}(4s^23d^m) = 17.58 + 0.95m \quad (m = 1-9)
$$
 (5)

and

$$
S_{5s}(5s^24d^m) = 35.11 + 0.94m \quad (m = 1-9)
$$
 (6)

Next we express the screening constant of 3d in  $4s<sup>2</sup>3d<sup>m+1</sup>$  as  $S_{3d}$  ( $4s<sup>1</sup>3d<sup>m+1</sup>$ ) or  $S_{3d}$  and that of 4d in  $5s^14d^{m+1}$  as  $S_{4d}(5s^14d^{m+1})$  or  $S_{4d}$ . The behaviors of  $S_{3d}$  and  $S_{4d}$  are more complicated than  $S_{4s}$  and  $S_{5s}$ . It is seen in Fig. 4 that the difference between  $S_{5s}$  and  $S_{4d}$  is much smaller than that between  $S_{4s}$  and  $S_{3d}$ . At the Mo, Pd, and Ag atoms,  $S_{4d}$  is almost equal or smaller than  $S_{5s}$ , reflecting a larger binding (i.e., IE) of the 4d electrons than the 5s electrons. In Fig. 5, we give the differences  $S_{3d} - S_{4s}$  and  $S_{4d} - S_{5s}$ . It is seen that  $S_{3d} - S_{4s}$  oscillates around 0.8 and the changes in  $S_{3d} - S_{4s}$  are approximately independent of the increase in the nuclear charge. On the other hand,  $S_{4d} - S_{5s}$  has a trend to decrease as the nuclear charge increases. We found that NHF  $S_{3d}$  and  $S_{4d}$  are roughly expressed as





 $\hat{\boldsymbol{\beta}}$ 



Fig. 5. Differences of screening constants between 4d and 5s and those between 3d and 4s given by  $ns^1(n-1)d^m$  core

and

$$
S_{4d}(5s^{1}4d^{m+1}) = S_{5s} + 0.70 - 0.08m \quad (m = 1-9). \tag{8}
$$

The 3d electron moves in the field of effective charge which is smaller than that of  $4s$ by a constant amount, while the 4d electron moves in the field of effective nuclear charge which increases faster than that for  $5s$  as the number of the  $d$  electrons increases. Since the 4d AO is more diffuse than the 3d AO, the screening effect for a 4d electron by the other 4d electrons is smaller than that for a 3d by the other *3ds;*  the screening for 4d increases more slowly than that for 3d, when the number of d electrons increases. Thus, the difference between  $S_{3d}$  and  $S_{4d}$  shown in Eqs. (7) and (8) is natural. (See also the line A-B in Fig. 5 which shows the difference of the Z dependence between  $S_{3d}$  and  $S_{4d}$  since that between  $S_{4s}$  and  $S_{5s}$  is almost the same.) The growing effective nuclear charge for 4d is expected to reverse the energetic order of  $5s^24\bar{d}^m$  and  $5s^14d^{m+1}$  at some value of m. We here note that the effective charge as well as the principal quantum number determines the ionization energies. From Eq. (4) one obtains the nuclear charge which reverses the energetic order of  $ns^{1}(n - 1)d^{m+1}$  and  $ns^{2}(n - 1)d^{m}$ :

$$
Z > nS_{(n-1)d} - (n-1)S_{ns}.
$$
 (9)

Using Eqs. (6), (8),  $m = Z - 38$ , and the relation (9), one may find that the energetic order reverses at an atom with  $Z = 39.3$  and it actually happens at the Zr atom  $(Z = 40)$  in Fig. 1. Using (5), (6),  $m = Z - 20$ , and (9), on the other hand, one may see that  $4s^13d^{m+1}$  never becomes lower than  $4s^23d^m$  so far as we consider the neutral first row transition metal atoms.

The screening constants for s electrons given by experimental IEs in Table 2 are described by

$$
S_{4s}(4s^23d^m: \text{exptl}) = 17.30 + 0.94m \quad (m = 1-9) \tag{10}
$$

and

$$
S_{5s}(5s^24d^m: \text{exptl}) = 34.59 + 0.94m \quad (m = 1-9), \tag{11}
$$

and those for d electrons are roughly expressed as

$$
S_{3d}(4s^{1}3d^{m+1}: \text{exptl}) = S_{4s} + 0.82 \quad (m = 1-9) \tag{12}
$$

Why do the second transition metal atoms prefer  $5s<sup>1</sup>4d<sup>m+1</sup>$  to  $5s<sup>2</sup>4d<sup>m</sup>$ ? 291

and

$$
S_{4d}(5s^14d^{m+1}: \text{exptl}) = S_{5s} + 1.02 - 0.09m \quad (m = 1-9). \tag{13}
$$

Using Eqs. (11), (13),  $m = Z - 38$ , and the relation (9), we find that the energetic order reverses at an atom with  $Z = 41.3$  and it actually happens at the Nb atom  $(Z = 41$ ; see [19]). The agreement between the NHF calculations and experiments is rather well. We thus confirmed the argument given by the NHF calculations.

In order to develop clear discussion, we have above fitted the screening constants given in Table 2 to a linear function of the number of electrons m. However, the screening constants in Table 2 can be directly used to derive the critical nuclear charge. Such results are discussed in the appendix.

Table 2 shows that the screening from the Ar core is almost complete but that from the Kr core is not. We thus summarize the reasons why the second row transition metal atoms prefer the configuration  $5s<sup>1</sup>4d<sup>m+1</sup>$ :

a) the incomplete screening of the Kr core, and

b) the smaller screening for 4d which is proportional to the number of  $4d$  electrons in the  $5s^14d^m$  core.

Concerning  $s^1 d^{m+1}$  and  $s^2 d^m$  configurations, we finally comment that the calculated ground state for Zr is  $5s<sup>1</sup>4d<sup>3</sup>5F$ , but  $5s<sup>2</sup>4d<sup>2</sup>3F$  is known to be the ground state experimentally [19]. This is only the case where the Hartree-Fock approximation fails to predict the true ground state among the atoms He through Xe.

#### *2.3. Atoms In through Xe*

From the atom In to Xe, the electronic configuration for the ground states is 5s<sup>2</sup>5p<sup>m</sup>. We compare the mean value of r,  $\langle r \rangle$ , of the outermost s and p orbitals of the atoms of the groups 1, 2, and 13–18 in Table 3. We see that  $\langle r \rangle_s$  and  $\langle r \rangle_p$  of the atoms in the same group increase as we move from the second to the fifth period with the sole exception of a small decrease in  $\langle r \rangle_s$  and  $\langle r \rangle_p$  at Ga compared with those of A1. We may briefly sum up the trend as

$$
\langle r \rangle_{5s} > \langle r \rangle_{4s} > \langle r \rangle_{3s} \gg \langle r \rangle_{2s}, \tag{14}
$$

$$
\langle r \rangle_{5p} > \langle r \rangle_{4p} > \langle r \rangle_{3p} \ge \langle r \rangle_{2p} \tag{15}
$$

and

$$
\langle r \rangle_{5p} - \langle r \rangle_{5s} > \langle r \rangle_{4p} - \langle r \rangle_{4s} > \langle r \rangle_{3p} - \langle r \rangle_{3s} \ge \langle r \rangle_{2p} - \langle r \rangle_{2s}.
$$
 (16)

In Table 4 we list the excitation energies of the lower excited states of the second to the fifth period atoms. It is seen that the excitation energies from  $ns^2np^m$  to  $n s<sup>1</sup> n p<sup>m+1</sup>$  of atoms in the groups 13 and 14 have a local maximum at the fourth period atoms as we move down the periodic table. In many cases, excitation energies of the remaining groups  $(1, 2, 15-18)$  decrease as the nuclear charge increases.

#### **3. Concluding remarks**

In this paper we developed numerical Hartree-Fock wave function from Rb to Xe. Total 332 states were calculated. We discussed why the second row transition metal

		Period					
Group	Atoms	2nd	3rd	4th	5th		
				$\langle r \rangle_{ns}$			
1	Li, Na, K, Rb	3.8737	4.2088	5.2437	5.6319		
$\overline{2}$	Be, Mg, Ca, Sr	2.6494	3.2529	4.2185	4.6330		
13	B, Al, Ga, In	1.9771	2.5993	2.4889	2.8445		
14	C. Si. Ge. Sn	1.5893	2.2071	2.2258	2.5859		
15	N. P. As. Sb	1.3323	1.9327	2.0297	2.3902		
16	O. S. Se. Te	1.1420	1.7207	1.8692	2.2280		
17	F. Cl. Br. I	1.0011	1.5556	1.7389	2.0944		
18	Ne, Ar, Kr, Xe	0.8921	1.4222	1.6294	1.9810		
				$\langle r \rangle_{np}$			
13	B. Al, Ga, In	2.2048	3.4339	3.4243	3.7780		
14	C. Si. Ge. Sn	1.7145	2.7522	2.8669	3.2483		
15	N. P. As. Sb	1.4096	2.3227	2.5123	2.9011		
16	O. S. Se. Te	1.2322	2.0607	2.2996	2.6914		
17	F. Cl. Br. I	1.0848	1.8420	2.1116	2.5018		
18	Ne, Ar, Kr, Xe	0.9653	1.6630	1.9516	2.3380		
		$\langle r \rangle_{np} - \langle r \rangle_{ns}$					
13	B, Al, Ga, In	0.2277	0.8346	0.9354	0.9335		
14	C, Si, Ge, Sn	0.1252	0.5451	0.6411	0.6624		
15	N, P, As, Sb	0.0773	0.3900	0.4826	0.5109		
16	O. S. Se. Te	0.0902	0.3400	0.4304	0.4634		
17	F. Cl. Br. I	0.0837	0.2864	0.3727	0.4074		
18	Ne, Ar, Kr, Xe	0.0732	0.2408	0.3222	0.3570		

**Table 3.** Mean value of r for the atoms with the ground state configuration of  $ns \cdot np^m$  (all in a.u.)

Table 4. Excitation energies from the ground configuration of  $ns<sup>1</sup> np<sup>m</sup>$  (all in a.u.)

			Period					
Group	Atoms	<b>States</b>		2nd	3rd	4th	5th	
1	Li, Na, K, Rb	ns(1)	${}^{2}S$	0.00000	0.00000	0.00000	0.00000	
		np(1)	$^{2}P$	0.06766	0.07250	0.05164	0.04737	
		$(n-1)d(1)$	$^{2}D$			0.08908	0.07740	
$\overline{2}$	Be, Mg, Ca, Sr	ns(2)	١s	0.00000	0.00000	0.00000	0.00000	
		ns(1)np(1)	3p	0.06152	0.06792	0.06243	0.03494	
			1 p	0.17829	0.14351	0.10185	0.09079	
		$ns(1)(n-1)d(1)$	$^3D$			0.08093	0.06115	
			'D			0.10724	0.09039	
13	B. Al. Ga. In	ns(2)np(1)	$^{2}P$	0.00000	0.00000	0.00000	0.00000	
		ns(1)np(2)	4p	0.07840	0.08584	0.11237	0.08582	
			$^{2}D$	0.21719	0.18502	0.21282	0.17677	
			${}^{2}S$	0.28097	0.23146	0.26013	0.22045	
			2P	0.35002	0.27063	0.29707	0.25251	





atoms prefer the configuration  $5s<sup>1</sup>4d<sup>m+1</sup>$ . The reasons are: (a) the incomplete screening of the Kr core and (b) the smaller screening for 4d which is proportional to the number of  $4d$  electrons in the  $5s<sup>1</sup>4d<sup>m</sup>$  core.

In atomic and molecular calculations, we often use analytical approximations of Hartree-Fock wave functions. The excitation energies, the mean distance of r given in this paper, the electron affinities, and the ionization energies given in



Fig. 6. Nuclear charge which gives  $ns^1(n-1)d^{m+1}$  stabler than  $ns^2(n-1)d^m$ . a) For Sc to Cu. b) For Y to Ag

the previous papers [9, 13-15] would be useful for constructing the basis sets for molecules including the transition metal atoms. Full details of the present results are available on 3.5 in diskettes upon request to TK or HT.

*Acknowledgement.* This work has been supported in part by a Grant-in-Aid for Scientific Research from the Ministry of Education, Science, and Culture of Japan.

#### **Appendix**

The critical nuclear charge specified by the inequality (9) can also be directly calculated from the screening constants given in Table 2. The results are depicted in Fig. 6. For the first transition metal atoms, the NHF and experimental screening constants give the same stable configurations with the sole exception of Ni which will be explained below; in most cases  $4s^23d^m$  is stabler than  $4s<sup>1</sup>3d<sup>m+1</sup>$ . The results confirm what we have discussed in the main text. One may wonder why Ni  $4s^23d^9$  is stabler than  $4s^23d^8$  in experiment although the experimental ground configuration is  $4s^23d^8$ . Moore [19] gives the ground state as  $4s^23d^{8}F_4$ . Even if one adopts the gravity of the energetic states generated from the same spatial and spin configuration with different total angular momentum, ordinary the ground configuration does not change. In Ni, however, the gravity of  $4s^23d^{8}$ <sup>3</sup> $F_4$ ,  ${}^3F_3$ , and  ${}^3F_2$  in 240 cm<sup>-1</sup> higher than that of  $4s<sup>1</sup>3d<sup>93</sup>D_3$ ,  $<sup>3</sup>D_2$ , and  $<sup>3</sup>D_1$ . So far as we discuss the electronic structure under the LS</sup></sup> coupling scheme, we may change the experimental ground configuration from  ${}^{3}F$ to  ${}^{3}D$ .

In the second row transition metal atoms, the calculated nuclear charge which satisfies the relation (9) always predicts that  $5s<sup>1</sup>4d<sup>m+1</sup>$  is stabler than  $5s<sup>2</sup>4d<sup>m</sup>$  except for Y and Tc (see Fig. 6b). Except for Tc, this is what we concluded in the main text based on a fitting of screening constants  $S$  to a linear function of number electrons m. We thus see good agreement between the NHF calculations and experiment, confirming again the discussion in the main text.

## **References**

- 1. Froese C (1966) J Chem Phys 45:1417
- 2. Froese-Fischer C (1972) At Data 4:301; (1973) At Data Nucl Data Tables 12:87
- 3. Mann JB (1973) At Data Nucl Data Tables 12:1
- 4. Fraga S, Karwowski J, Saxena KM (1976) Handbook of atomic data. Elsevier, Amsterdam
- 5. Froese-Fischer C (1977) The Hartree-Fock method for atoms. Wiley, New York
- 6. Bunge CF, Barrientos JA, Bunge AV, Cogordan JA (1992) Phys Rev A46:3691
- 7. Partridge H (1987) NASA Technical Memorandum 89449, Ames Research Center, Moffett Field, CA
- 8. Partridge H (1989) NASA Technical Memorandum 101044, Ames Research Center, Moffett Field, CA
- 9. Koga T, Tatewaki H, Thakkar AJ (1994) J Chern Phys 100:8140
- 10. Hartree DR, Hartree W (1936) Proc Roy Soc (London) A154:588
- 11. Froese C (1968) J Chem Phys 47:4010
- 12. Hay PJ (1977) J Chem Phys 66:4377
- 13. Tatewaki H, Koga T, Sakai Y, Thakkar AJ (1994) J Chem Phys 101:4945
- 14. Tatewaki H, Koga T (1994) Chem Phys Lett 228:562
- 15. Tatewaki H, Sekiya M, Sasaki F, Matsuoka O, Koga T (1995) Phys Rev A51:197
- 16. Froese-Fischer C (1972) Comput Phys Commun 4:107
- 17. Tatewaki H, Taketa K, Sasaki F (1971) lnt J Quantum Chem 5:335
- 18. Tatewaki H, Tanaka K (1974) J Chem Phys 60:601
- 19. Moore CE (1971) Atomic Energy Levels, NSRDS-NBS 35, Vols 1-3, Nat. Bur. Stand. US, Washington